# **PHOTONICS** Research

# Snapshot spectral compressive imaging reconstruction using convolution and contextual Transformer

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Spectral compressive imaging (SCI) is able to encode a high-dimensional hyperspectral image into a twodimensional snapshot measurement, and then use algorithms to reconstruct the spatio-spectral data-cube. At present, the main bottleneck of SCI is the reconstruction algorithm, and state-of-the-art (SOTA) reconstruction methods generally face problems of long reconstruction times and/or poor detail recovery. In this paper, we propose a hybrid network module, namely, a convolution and contextual Transformer (CCoT) block, that can simultaneously acquire the inductive bias ability of convolution and the powerful modeling ability of Transformer, which is conducive to improving the quality of reconstruction to restore fine details. We integrate the proposed CCoT block into a physics-driven deep unfolding framework based on the generalized alternating projection (GAP) algorithm, and further propose the GAP-CCoT network. Finally, we apply the GAP-CCoT algorithm to SCI reconstruction. Through experiments on a large amount of synthetic data and real data, our proposed model achieves higher reconstruction quality (>2 dB in peak signal-to-noise ratio on simulated benchmark datasets) and a shorter running time than existing SOTA algorithms by a large margin. The code and models are publicly available at https://github.com/ucaswangls/GAP-CCoT. © 2022 Chinese Laser Press

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# **1. INTRODUCTION**

A hyperspectral image is a spatio-spectral data-cube consisting of many narrow spectral bands, each spectral band corresponding to a wavelength. Compared with RGB images, hyperspectral images have rich spectral information and can be widely used in medical diagnosis [1], food safety [2], remote sensing [3], and other fields. However, the long imaging time and high hardware cost of existing hyperspectral cameras greatly limit the application of these devices. To address the above problems, spectral compressive imaging (SCI), especially the coded aperture snapshot spectral imaging (CASSI) system [1,4,5], provides an elegant solution, which can capture information from multiple spectral bands at the same time with only one twodimensional (2D) sensor. CASSI uses a physical mask and a prism to modulate the spectral data-cube, and captures the modulated and compressed measurement on a 2D plane sensor. Then reconstruction algorithms are employed to recover the hyperspectral data-cube from the measurement along with the mask. This paper focuses on the reconstruction algorithm.

At present, SCI reconstruction algorithms mainly include model-based methods and learning-based methods. Traditional model-based methods have relevant theoretical proofs and can be well explained. The representative algorithms are mainly TWo-step Iterative Shrinkage/Thresholding algorithm (TwIST) [6], generalized alternating projection total variation (GAP-TV) [7], and DEcompress SCI (DeSCI) [8]. However, model-based methods require prior knowledge and long reconstruction times and usually provide only poor reconstruction quality. With its strong fitting ability, a deep learning model can directly learn the relevant knowledge from data and provide excellent reconstruction results [9–13]. However, compared to model-based methods, learning-based methods lack interpretability [14].

The deep unfolding network driven by physics combines the advantages of model-based and learning-based methods, so it is powerful with clear interpretability [15–18]. At present, most advanced reconstruction algorithms [19,20] are based on the idea of deep unfolding. Many models combine U-net [21] with the deep unfolding idea for image reconstruction and achieve good reconstruction results. However, the U-net model is too simple to fully capture the effective information of the image. Therefore, we use the inductive bias ability of



**Fig. 1.** Reconstructed real data of Legoman, captured by snapshot SCI systems in Ref. [20]. We show reconstruction results of 12 spectral channels, and compare our proposed method with the latest self-supervised method (PnP-DIP-HSI [23]) and the method based on maximum *a posteriori* (MAP) estimation (DGSMP algorithm [24]). As can be seen from the purple and green areas in the plot, our method reconstructs a clearer image, the PnP-DIP-HSI method produces some artifacts, and the DGSMP method loses some details.

convolution and the powerful modeling ability of Transformer [22] to design a parallel module to solve the problem of SCI reconstruction. As shown in Fig. 1, the integration of our proposed method and deep unfolding idea can recover more details with fewer artifacts.

Our main contributions in this paper are summarized as follows:

• we first apply Transformer to deep unfolding for SCI reconstruction;

• we propose an effective parallel network structure composed of convolution and contextual Transformer (CCoT), which can obtain more spectral features;

• experimental results on a large amount of synthetic and real data show that our proposed method achieves state-of-theart (SOTA) results in SCI reconstruction;

• the proposed method can also be used in other compressive sensing (CS) systems [25,26], such as video CS [27–29], and yields excellent results.

# 2. RELATED WORK

In this section, we first review the forward model of CASSI, and then briefly introduce existing reconstruction methods. Focusing on deep-learning-based models, we describe the pros and cons of convolutional neural networks (CNNs) and introduce the vision Transformer (ViT) for other tasks.

#### A. Mathematical Model of SCI System

The SCI system encodes a high-dimensional spectral data-cube into 2D measurement, and CASSI [4] is one of the earliest SCI systems. As shown in Fig. 2, the three-dimensional (3D) spatiospectral data-cube is first modulated by a coded aperture (a.k.a., mask). Then, the encoded 3D spectral data-cube is dispersed by the prism. Finally, the entire (modulated) spectral data-cube is captured by a 2D camera sensor by integrating across the spectral dimension.

Let  $F \in \mathbb{R}^{n_x \times n_y \times n_\lambda}$  denote the captured 3D spectral datacube, and  $M \in \mathbb{R}^{n_x \times n_y}$  denote a pre-defined mask, where  $n_x$ ,  $n_y$ , and  $n_\lambda$  represent the height, width, and channel number of the spectral image, respectively. For each spectral channel  $\ell' = 1, ..., n_\lambda$ , the spectral image is modulated to  $F'_{\ell} = F_{\ell} \odot M$ , where  $F_{\ell}$  and  $F'_{\ell}$  represent the original and modulated spectral images at the  $\ell$ th spectral channel, respectively, and  $\odot$  denotes element-wise multiplication. Then after passing through the dispersive prism, the modulated spectral data-cube is tilted. Finally, by compressing across the spectral domain, the camera sensor captures a 2D compressed measurement  $G \in \mathbb{R}^{n_x \times (n_y + n_z - 1)}$ , which can be expressed as

$$G_{u,v} = \sum_{\ell=1}^{n_{\lambda}} F_{u,v,\ell}^{\prime\prime} + Z_{u,v}, \qquad (1)$$

where (u, v) represents the coordinate system of the camera detector plane,  $F''_{u,v,l} = F'_{x,y+d(\lambda_\ell - \lambda_c),\ell}$  denotes the tilted spectral data-cube after passing through the dispersive prism of the  $\ell$ th spectral channel, (x, y) represents the coordinate system of each modulated spectral image,  $d(\lambda_\ell - \lambda_c)$  represents the spatial shifting of the  $\ell$ th spectral channel where d is a scalar,  $\lambda_\ell$  is the wavelength at the  $\ell$ th channel,  $\lambda_c$  denotes the reference wavelength that does not shift after passing through the disperser, and  $Z \in \mathbb{R}^{n_x \times (n_y + n_{\lambda} - 1)}$  denotes the measurement noise.

For the sake of simple notations, as derived in Ref. [23], we further give the vectorized formulation expression of Eq. (1). First, we define  $\operatorname{vec}(\cdot)$  as a vectorization operation of a matrix. Then we vectorize  $\boldsymbol{g} = \operatorname{vec}(\boldsymbol{G}) \in \mathbb{R}^{n_x(n_y+n_\lambda-1)}, \ \boldsymbol{z} = \operatorname{vec}(\boldsymbol{Z}) \in \mathbb{R}^{n_x(n_y+n_\lambda-1)}, \ \text{and} \ \boldsymbol{f} = [\boldsymbol{f}_1^T, \dots, \boldsymbol{f}_{n_\lambda}^T]^T \in \mathbb{R}^{n_x n_y n_\lambda},$  where  $\boldsymbol{f}_{\ell} = \operatorname{vec}(\boldsymbol{F}_{\ell})$ . In addition, we define the sensing matrix generated by a coded aperture and disperser in the CASSI system as

$$H = [\boldsymbol{D}_1, \dots, \boldsymbol{D}_{n_{\lambda}}] \in \mathbb{R}^{n_x(n_y + n_{\lambda} - 1) \times n_x n_y n_{\lambda}},$$
(2)

where  $\boldsymbol{D}_{\ell} = \begin{bmatrix} \mathbf{0}^{(1)} \\ A_{\ell} \\ \mathbf{0}^{(2)} \end{bmatrix} \in \mathbb{R}^{n_x(n_y + n_{\lambda} - 1) \times n_x n_y}$ , where  $\boldsymbol{A}_{\ell} =$ 

 $\begin{array}{l} \text{Diag}(\text{vec}(\pmb{M})) \in \mathbb{R}^{n_x n_y \times n_x n_y} \text{ is a diagonal matrix and its} \\ \text{diagonal element is } \text{vec}(\pmb{M}), \text{ and } \pmb{0}^{(1)} \in \mathbb{R}^{(\ell-1) \times n_x n_y} \text{ and} \end{array}$ 



Fig. 2. Schematic diagrams of CASSI system.

 $\mathbf{0}^{(2)} \in \mathbb{R}^{(n_{\lambda}-\ell) \times n_{x}n_{y}}$  represent the zero matrix. Finally, the vectorization expression of Eq. (1) is

$$g = Hf + z. \tag{3}$$

After obtaining the measurement g, the next task is to develop a decoding algorithm. Given g and H, solve f.

## **B.** Reconstruction Algorithms for SCI

SCI reconstruction algorithms mainly focus on how to solve the ill-posed inverse problem in Eq. (3), a.k.a., the reconstruction of SCI. Traditional methods are generally based on prior knowledge as a regularization condition to solve the problem, such as using TV [6], sparsity [30], dictionary learning [31,32], non-local low rank [8,33], and Gaussian mixture modes [34]. The main problem of these algorithms is that they need to manually set prior knowledge and iteratively solve the problem. Therefore, the reconstruction time is long, and the quality is usually not good.

With its powerful learning capability, the neural network can directly learn a mapping relationship from the measurement to the original hyperspectral image, and the reconstruction speed can reach the millisecond level. End-to-end (E2E) deep learning methods (Spatial-Spectral Self-Attention network (TSA-net) [35],  $\lambda$ -net [9], Spatial/Spectral Invariant Residual U-Net (SSI-ResU-Net) [10]) take the measurement and masks as inputs, and use only one single network to reconstruct the desired signal directly. Plug-and-play (PnP) methods [36,37] use a pre-trained network as a denoiser plugged into iterative optimization [7,38]. Different from PnP methods, the denoising networks at each stage of deep unfolding methods [19,20] are independent of each other, and the parameters are not shared, and thus can be trained E2E.

Deep unfolding is driven by physics and offers the advantages of high-speed, high-quality reconstruction while enjoying the benefits of physics-driven interpretability. Therefore, in this paper, we follow the deep unfolding framework [20] and propose a new deep denoiser block based on CCoT. The proposed module along with deep unfolding leads to SOTA results for SCI reconstruction.

#### **C. Limitations of CNNs for Reconstruction**

Due to local connection and shift-invariance, the convolutional network [39] can well extract local features of images, and is widely used in image recognition [40–42], object detection [43], semantic segmentation [44], image denoising [45], and other tasks [46,47]. However, its local connection property also makes it lack the ability of global perception. To improve the receptive field of convolution, deeper network architecture [41] or various pooling operations [48] are often used. The squeeze-and-excitation network (SENet) [48] uses the channel attention (CA) mechanism [49] to aggregate the global context and redistributes the weight to each channel. However, these methods usually lose a significant amount of detail information and are not friendly to image reconstruction and other tasks that need to recover local details.

Bearing the above concerns and considering the running time, we do not use very deep network structure in our work for SCI reconstruction. Instead, we use a convolution with a sliding step of two instead of the traditional max pooling operation, aiming to capture the local details of the desired spatio-spectral data-cube.

ViT [50] and its variants [51-54] have verified the effectiveness of Transformer architecture in computer vision tasks. However, training a good ViT model requires a large number of training datasets (i.e., JFT-300M [55]), and its computational complexity increases quadratically with image size. To better apply Transformer to computer vision related tasks, the latest Swin Transformer [56] proposes a local window self-attention mechanism and a shifting window method, which greatly reduces computational complexity. The Transformer network based on Swin has achieved amazing results in computer vision tasks such as image recognition [57], object detection [58], semantic segmentation [59,60], and image restoration [61], which further verifies the feasibility of Transformer in computer vision. In addition, when computing self-attention, most Transformers including Swin Transformer are independently learned for all pairwise query-keys, without using the rich contextual relations between them. Moreover, the self-attention mechanism in ViTs often ignores local feature details, which is not conducive to low-level image tasks such as image reconstruction.

Inspired by contextual Transformer (CoT) [62] and conformer networks [63], in this paper, we propose a network structure named CCoT, which can take advantage of convolution and Transformer to extract more effective spectral features, and can be well applied to image reconstruction tasks such as SCI.

# **3. PROPOSED NETWORK**

In this section, we first briefly review the GAP-net [20] algorithm, which uses deep unfolding ideas [64] and the GAP algorithm [65] for SCI reconstruction. We select GAP-net because of its high performance, robustness, and flexibility for different SCI systems reported in Ref. [20]. Following this, we combine the advantages of convolution and Transformer and then propose a module named CCoT. We integrate this module into GAP-net to reconstruct hyperspectral images from the compressed measurements and masks.

#### A. Review of GAP-net for SCI Reconstruction

The SCI reconstruction algorithm is used to solve the following optimization problem:

$$\hat{f} = \arg\min_{f} \frac{1}{2} \|\boldsymbol{g} - \boldsymbol{H}\boldsymbol{f}\|^{2} + \lambda \Omega(\boldsymbol{f}), \qquad (4)$$

where the first term is the fidelity term, and the second term,  $\Omega(f)$ , is the prior or regularization to confine the solutions. In GAP-net and other deep unfolding algorithms, implicit priors (represented by deep neural networks) have been used to improve the performance.

Following the framework of GAP, Eq. (4) can be rewritten as a constrained optimization problem by introducing an auxiliary parameter v:

$$(\hat{f}, \hat{\boldsymbol{v}}) = \operatorname*{arg\,min}_{f, \boldsymbol{v}} \frac{1}{2} \|\boldsymbol{f} - \boldsymbol{v}\|_2^2 + \lambda \Omega(\boldsymbol{v}), \quad \text{s.t.} \quad \boldsymbol{g} = H\boldsymbol{f}.$$
 (5)

To solve Eq. (5), GAP decomposes it into the following subproblems for iterative solutions, with  $\eta$  denoting the iteration number.



**Fig. 3.** Architecture of the proposed GAP-CCoT. (a) GAP-net with *N* stages;  $\mathcal{G}(\cdot)$  represents the operation of Eq. (6),  $\mathcal{D}(\cdot)$  represents a denoiser, and  $\boldsymbol{v}^{(0)} = \boldsymbol{H}^T \boldsymbol{g}$ . (b) CCoT-net, the proposed denoising network plugged into GAP algorithm. (c) Convolution branch and Transformer branch; the output is connected with concatenation. (d) Convolution block with channel attention; *c* represents the output number of convolution channels. (e) Contextual Transformer block. (f) Pixelshuffle algorithm for fast upsampling.

• Solving  $f: f^{(\eta+1)}$  is updated via an Euclidean projection of  $v^{(\eta)}$  on the linear manifold  $\mathcal{M}: g = Hf$ :

$$f^{(\eta+1)} = v^{(\eta)} + H^T (HH^T)^{-1} (g - Hv^{(\eta)}).$$
 (6)

• Solving v: we can apply a trained denoiser to map f closer to the desired signal space:

$$v^{(\eta+1)} = \mathcal{D}_{\eta+1}(f^{(\eta+1)}),$$
 (7)

where  $\mathcal{D}_{\eta+1}$  denotes the denoising operation.

It has been derived in the literature [7] that Eq. (6) has a closed-form solution due to the special structure of H in Eq. (2). Therefore, the only difference (and novelty) is the denoising step in Eq. (7). In the following, we describe the novel CCoT block proposed in this work for efficient and effective SCI reconstruction. The general reconstruction framework is illustrated in Fig. 3(a), and the detailed CCoT block is depicted in Figs. 3(b)–3(f).

#### **B. Proposed CCoT Block for Deep Denoising**

As mentioned in Section 2.D, to address the challenge of SCI reconstruction, we develop the CCoT block, in which convolution and Transformer are used in parallel, which can be well applied to image reconstruction tasks such as SCI.

### 1. Convolution Branch

As shown in Figs. 3(c) and 3(d), the convolution branch consists of a down-sampling layer and a CA block. In this paper, we use convolution layer for down-sampling with sliding step *s* instead of direct max pooling to capture fine details, and s = 2 is used in the experiments. The CA block draws lessons from the idea of SENet [48], automatically obtains the importance of each feature channel by learning, then improves useful features according to this importance and suppresses features that are not significant for the current task. The first convolution layer and CA module are followed by a LeakyReLU activation function [66]. The proposed convolution branch can extract local features of images well.

#### 2. Contextual Transformer Branch

By calculating the similarity between pixels, the traditional Transformer makes the model focus on different regions and extract more effective features. However, when calculating paired query-keys, they are relatively independent. A single spectral image itself contains rich contextual information, and there is also a significant amount of correlations between adjacent spectra. Therefore, we designed a CoT branch to better obtain features of hyperspectral images.

As shown in Fig. 3(c), the CoT branch consists of a downsampling layer and a CoT block. The structure of the downsampling layer is the same as the convolution branch. As shown in Fig. 3(e), we first recall that the input of the hyperspectral image is of  $F \in \mathbb{R}^{n_x \times n_y \times n_\lambda}$ . Then we define queries, keys, and values as  $K_1 \in \mathbb{R}^{n_x \times n_y \times n_\lambda}$ ,  $Q \in \mathbb{R}^{n_x \times n_y \times n_\lambda}$ ,  $V \in \mathbb{R}^{n_x \times n_y \times n_\lambda}$ , respectively. Different from the traditional self-attention that uses  $1 \times 1$  convolutions to generate mutually independent paired query-keys, the CoT block first applies the group convolution of size  $m \times m$  to generate a static key  $K_1 \in \mathbb{R}^{n_x \times n_y \times n_\lambda}$  containing the context, and  $K_1$  can be used as a static context representation of input F. Q and V can be generated by the traditional self-attention mechanism. Then, we concatenate  $K_1$  and Q by the third dimension (spectral channels), followed by two  $1 \times 1$ convolutions to generate an attention matrix:

$$\boldsymbol{A} = \operatorname{Conv}_1(\operatorname{Conv}_2([\boldsymbol{K}_1, \boldsymbol{Q}]_3)), \quad (8)$$

where  $[]_3$  denotes the concatenation along the third dimension,  $\text{Conv}_1, \text{Conv}_2$  represent two  $1 \times 1$  convolutions,  $A \in \mathbb{R}^{n_x \times n_y \times (m^2 \times C_b)}$  represents the attention matrix containing context, and  $C_b$  denotes the number of attention heads. We use the traditional self-attention mechanism to perform a weighted summation of V through A to obtain the dynamic context  $K_2 \in \mathbb{R}^{n_x \times n_y \times n_\lambda}$ , and then fuse dynamic context  $K_2$ and static context  $K_1$  as the output of the CoT block through the attention mechanism [48].

Finally, we concatenate the output of the convolution branch and CoT branch as the final output of the CCoT block.

# C. GAP-CCoT Network

As shown in Fig. 3(b), we use the CCoT module and pixelshuffle algorithm to construct a U-net [21] like network as the denoiser in GAP-net. The network consists of a contracting path and an expansive path. The contracting path contains three CCoT modules, and the expansive path contains three up-sampling modules. Each module of the expansive path is first quickly up-sampled by the pixelshuffle algorithm [67], followed by a  $3 \times 3$  convolution, and finally concatenates the output from the corresponding stage of the contracting path (after a  $1 \times 1$  convolution) as the input of the next module. Eventually, CCoT, GAP, and deep unfolding form the reconstruction network (GAP-CCoT) of SCI.

Last, following GAP-net [20] and hyperspectral image reconstruction using a deep Spatial-Spectral Prior (HSSP) [19] network, the loss function of the proposed model is

$$\mathcal{L}_{\text{MSE}}(\Theta) = \frac{1}{n_{\lambda}} \sum_{\ell=1}^{n_{\lambda}} \|\hat{F}_{\ell} - F_{\ell}\|_{2}^{2}, \qquad (9)$$

where  $\mathcal{L}_{MSE}(\Theta)$  represents the mean square error (MSE) loss,  $n_{\lambda}$  again represents the spectral channel to be reconstructed,

and  $\hat{F}_{\ell} \in \mathbb{R}^{n_x \times n_y}$  is the reconstructed hyperspectral image at the  $\ell$ th spectral channel.

# 4. EXPERIMENTAL RESULTS

In this section, we compare the performance of the proposed GAP-CCoT network with several SOTA methods on both simulation and real datasets. The peak-signal-to-noise-ratio (PSNR) and structured similarity index metrics (SSIM) [68] are used to evaluate the performance of different hyperspectral image reconstruction methods.

# A. Datasets

We use the hyperspectral dataset CAVE [69] for model training and KAIST [70] for model simulation testing. The CAVE dataset consists of 32 scenes, including full spectral resolution reflectance data from 400 nm to 700 nm with a 10 nm step, and a spatial resolution of  $512 \times 512$ . The KAIST dataset consists of 30 scenes with a spatial resolution of  $2704 \times 3376$ . To match the wavelength of the real CASSI system, we follow the method proposed by TSA-net [71] and employ the spectral interpolation method to modify the training set and test data wavelength. The final wavelength was fitted to 28 spectral bands ranging from 450 nm to 650 nm.

# **B.** Implementation Details

During training, we use random cropping, rotation, and flipping for CAVE dataset augmentation. By simulating the imaging process of CASSI, we can obtain the corresponding measurement. We use measurement and masks as inputs to train GAP-CCoT and use the Adam optimizer [72] to optimize the model. The learning rate is set to 0.001 initially and reduces by 10% every 10 epochs. Our model is trained for 200 epochs in total. All experiments are run on the NVIDIA RTX 8000 GPU using PyTorch.

Finally, we use a GAP-CCoT network with nine stages as the reconstruction network, and no noise is added to the measurement during training on simulation data. We added shot noise to the measurements for model training on real data following the procedure in Ref. [20].

# **C. Simulation Results**

We compare the method proposed in this paper with several SOTA methods (TwIST [6], GAP-TV [7], DeSCI [8], HSSP [19],  $\lambda$ -net [9], TSA-net [71], GAP-net [20], Plug-and-Play Deep Image Priors Hyperspectral Images (PnP-DIP-HSI) [23], Deep Gaussian Scale Mixture Prior (DGSMP) [24] and SSI-ResU-Net (v1) [10]) on synthetic datasets. Table 1 presents the average PSNR and SSIM results of different spectral reconstruction algorithms. We can see that the average PSNR value of our proposed algorithm is 35.26 dB and average SSIM value is 0.950. The average PSNR value is 2.09 dB higher than that of the current best algorithm SSI-ResU-Net (v1, preprinted, not published), and the SSIM value is 0.021 higher. In addition, compared with the self-supervised learning method PnP-DIP-HSI and DGSMP method (best published results) based on the maximum a posteriori (MAP) estimation, the average PSNR of our proposed method is 3.96 dB and 2.63 dB higher, respectively. Based on these significant improvements, we can conclude the powerful learning capability of Transformer and the proposed CCoT block.

Table 1. Average PSNR in dB (upper entry in each cell) and SSIM (lower entry in each cell) of Different Algorithms on 10 Synthetic Datasets<sup>a</sup>

Algorithms	Scene 1	Scene 2	Scene 3	Scene 4	Scene 5	Scene 6	Scene 7	Scene 8	Scene 9	Scene 10	Average
TwIST [6]	24.81	19.99	21.14	30.30	21.68	22.16	17.71	22.39	21.43	22.87	$22.44 \pm 3.32$
	0.730	0.632	0.764	0.874	0.688	0.660	0.694	0.682	0.729	0.595	$0.704 \pm 0.077$
GAP-TV [7]	25.13	20.67	23.19	35.13	22.31	22.90	17.98	23.00	23.36	23.70	$23.73 \pm 4.45$
	0.724	0.630	0.757	0.870	0.674	0.635	0.670	0.624	0.717	0.551	$0.685 \pm 0.088$
DeSCI [8]	27.15	22.26	26.56	39.00	24.80	23.55	20.03	20.29	23.98	25.94	$25.35\pm5.38$
	0.794	0.694	0.877	0.965	0.778	0.753	0.772	0.740	0.818	0.666	$0.785 \pm 0.087$
HSSP [19]	31.48	31.09	28.96	34.56	28.53	30.83	28.71	30.09	30.43	28.78	$30.35 \pm 3.79$
	0.858	0.842	0.832	0.902	0.808	0.877	0.824	0.881	0.868	0.842	$0.852 \pm 0.049$
$\lambda$ -net [9]	30.82	26.30	29.42	36.27	27.84	30.69	24.20	28.86	29.32	27.66	$29.14 \pm 3.20$
	0.880	0.846	0.916	0.962	0.866	0.886	0.875	0.880	0.902	0.843	$0.886 \pm 0.035$
TSA-net [71]	31.26	26.88	30.03	39.90	28.89	31.30	25.16	29.69	30.03	28.32	$30.15 \pm 3.92$
	0.887	0.855	0.921	0.964	0.878	0.895	0.887	0.887	0.903	0.848	$0.893 \pm 0.033$
PnP-DIP-HSI [23]	32.70	27.27	31.32	40.79	29.81	30.41	28.18	29.45	34.55	28.52	$31.30 \pm 3.98$
	0.898	0.832	0.920	0.970	0.903	0.890	0.913	0.885	0.932	0.863	$0.901 \pm 0.038$
GAP-net [20]	33.03	29.52	33.04	41.59	30.95	32.88	27.60	30.17	32.74	29.73	$32.13 \pm 3.81$
	0.921	0.903	0.940	0.972	0.924	0.927	0.921	0.904	0.927	0.901	$0.924\pm0.021$
DGSMP [24]	33.26	32.09	33.06	40.54	28.86	33.08	30.74	31.55	31.66	31.44	$32.63 \pm 3.07$
	0.915	0.898	0.925	0.964	0.882	0.937	0.886	0.923	0.911	0.925	$0.917 \pm 0.024$
SSI-ResU-Net (v1) [10]	34.06	30.85	33.14	40.79	31.57	34.99	27.93	33.24	33.58	31.55	$33.17 \pm 3.34$
	0.926	0.902	0.924	0.970	0.939	0.955	0.861	0.949	0.931	0.934	$0.929 \pm 0.030$
Ours	35.17	35.90	36.91	42.25	32.61	34.95	33.46	33.13	35.75	32.43	$35.26 \pm 2.89$
	0.938	0.948	0.958	<b>0.9</b> 77	0.948	<b>0.95</b> 7	0.923	0.952	0.954	0.941	$0.950 \pm 0.014$

"Best results are in bold.

Figure 4 shows part of the visualization results and spectral curves of two scenes using several SOTA spectral SCI reconstruction algorithms. Enlarging the local area, we can see that our proposed method can recover more edge details and better spectral correlation than other algorithms.

In addition, we also analyze the computational complexity of our method and compare it with several previous deeplearning-based SOTA spectral reconstruction algorithms. As shown in Table 2, our proposed GAP-CCoT-S3 (with three stages) achieves higher reconstruction quality than previous SOTA algorithms with lower computational cost.

#### D. Flexibility of GAP-CCoT to Mask Modulation

CCoT-net serves only as a denoiser for the GAP algorithm, so the GAP-CCoT network proposed in this paper has flexibility for different signal modulations. To verify this, we train the GAP-CCoT network on one mask and test it on five other untrained masks. Table 3 shows the test results of the average PSNR value and SSIM value on 10 simulation data using different masks (five new masks of size  $256 \times 256$  randomly cropped from the real mask of size  $660 \times 660$ ). We can observe that for a new mask that does not appear in training, the average PSNR decrease remains within 0.27 dB, which is still better



Fig. 4. Reconstruction results of GAP-CCoT and other spectral reconstruction algorithms ( $\lambda$ -net, HSSP, TSA-net, GAP-net, DGSMP, PnP-DIP-HSI) in scene 3 and scene 9. Zoom in for better view.

Table 2.Computational Complexity and AverageReconstruction Quality of Several SOTA Algorithms on10 Synthetic Datasets

Algorithm	Params (10 <sup>6</sup> )	FLOPs (10 <sup>9</sup> )	PSNR (dB)	SSIM
λ-net [9]	66.16	514.33	29.25	0.886
TSA-net [71]	44.25	135.03	30.15	0.893
GAP-net [20]	2.89	54.16	32.13	0.924
DGSMP [24]	3.76	647.28	32.63	0.917
SSI-ResU-Net (v1) [10]	1.25	81.98	33.17	0.929
GAP-CCoT-S3	2.68	31.84	33.89	0.934
GAP-CCoT-S9	8.04	95.52	35.26	0.950

Table 3.Average PSNR and SSIM Results on 10Synthetic Data with Different Masks

Mask	PSNR (dB)	SSIM
Mask used in training	$35.26 \pm 2.89$	$0.950 \pm 0.014$
New mask 1	$35.10 \pm 2.92$	$0.949 \pm 0.015$
New mask 2	$35.06 \pm 2.91$	$0.948 \pm 0.015$
New mask 3	$35.06 \pm 2.91$	$0.949 \pm 0.015$
New mask 4	$35.02 \pm 2.92$	$0.948\pm0.014$
New mask 5	$34.99 \pm 2.90$	$0.948\pm0.014$



**Fig. 5.** Architecture of the proposed Stacked CCoT. The input of the network is  $H^T g$ , and CCoT-net is the same as in Fig. 3(b).

than other algorithms. Therefore, we can conclude that the GAP-CCoT network proposed in this paper is flexible for large-scale SCI reconstruction.

# E. Ablation Study

To verify the effectiveness of CoT and GAP algorithms, we trained two different GAP-CCoT networks and two different Stacked CCoT networks (shown in Fig. 5) for spectral SCI reconstruction, respectively. Table 4 shows the reconstruction results of the proposed two networks, where "w/o" CoT means removing the CoT branch at each stage of coding. We can clearly observe that the GAP-CCoT network is 0.99 dB higher in PSNR than the Stacked CCoT network. The PSNR value of the CoT module is improved by 1.13 dB and 1.41 dB on the GAP-CCoT network, respectively.

 Table 4.
 Ablation Study: Average PSNR and SSIM Values

 of Different Algorithms on 10 Synthetic Data

Algorithms	PSNR (dB)	SSIM
Stacked CCoT w/o CoT	$32.86 \pm 3.01$	$0.924 \pm 0.021$
GAP-CCoT w/o CoT	$34.13 \pm 2.95$	$0.933 \pm 0.019$
Stacked CCoT	$34.27 \pm 2.94$	$0.936 \pm 0.018$
GAP-CCoT	$35.26\pm2.89$	$0.950\pm0.014$





Table 5.Computational Complexity and AverageReconstruction Quality of GAP-CCoT on 10 SyntheticData with Different Stages

Stage Number	Params (10 <sup>6</sup> )	FLOPs (10 <sup>9</sup> )	PSNR (dB)	SSIM
3	2.68	31.84	33.89	0.934
5	4.47	53.06	34.30	0.936
7	6.25	74.29	34.86	0.940
9	8.04	95.52	35.26	0.950
12	10.72	127.35	35.43	0.951
15	13.41	159.19	35.54	0.952

Table 6. Average PSNR and SSIM Results on 10Synthetic Data with Different Loss Functions

Loss Function	PSNR (dB)	SSIM	
LAD	35.48	0.952	
MSE	35.26	0.950	

To verify the impact of the number of stages on the reconstruction quality, we trained multiple models with different numbers of stages. As can be seen from Fig. 6 and Table 5, the model proposed in this paper needs only three stages to achieve high reconstruction quality, and the reconstruction quality improves with the increase in number of stages, but the computational complexity also increases. In addition, we also notice that the spectral reconstruction quality improves slowly after nine stages. To trade off between accuracy and computational complexity, we set the number of stages to nine.

To verify the effect of the loss function on reconstruction quality, we use the least absolute deviation (LAD) loss function to retrain our proposed model. As shown in Table 6, our method can further improve the reconstruction quality by using the LAD loss function.

#### F. Real Data Results

We test the proposed method on several real data captured by the CASSI system [4,71]. The system captures 28 spectral bands with wavelengths ranging from 450 nm to 650 nm. The spatial resolution of the object is  $550 \times 550$ , and the spatial resolution of the measurements captured by the plane sensor is  $550 \times 604$ . Due to the flexibility of our proposed method for



**Fig. 7.** Reconstruction results of GAP-CCoT and other spectral reconstruction algorithms ( $\lambda$ -net, TSA-net, GAP-net, DGSMP, PnP-DIP-HSI) in two real scenes (scene 1 and scene 2).

different masks, we trained GAP-CCoT with a mask of spatial size  $256 \times 256$  and directly applied it to real measurements with a spatial size of  $550 \times 640$ . We compared our method with several SOTA methods ( $\lambda$ -net [9], TSA-net [71], GAP-net [20], PnP-DIP-HSI [23], DGSMP [24]) on real data. In addition to the results shown in Fig. 1, Fig. 7 shows partial visualization reconstructed results and spectral curves of real data from another scene. By zooming in on a local area, we can see that our proposed method can recover more details and has fewer artifacts. In addition, from the spectral correlation curve, our proposed method also achieves higher spectral accuracy than existing methods.

# 5. CONCLUSION AND DISCUSSION

In this paper, we use the inductive bias ability of convolution and the powerful modeling ability of Transformer to propose a parallel module, named CCoT, which can obtain more effective spectral features. We integrate this module with a physicsdriven deep unfolding idea and GAP algorithm, which can be well applied to SCI reconstruction.

In addition, we have also developed similar models for video CS [14,27,73] and our model produces excellent results, which are summarized in Table 7 and Fig. 8. We can see that our

method can achieve higher reconstruction quality and more details. As shown in Table 8, we further analyze the computational complexity of GAP-CCoT and compare it with previous SOTA reconstruction algorithms. Due to the addition of the CA mechanism and the Transformer module, our algorithm has more parameters and running time than some previous deeplearning-based algorithms (U-net, MetaSCI, GAP-net), but these modules bring about a significant improvement in reconstruction quality, and our proposed method maintains a high real-time performance (0.064 s). Moreover, it has better real-time performance than other high-precision reconstruction algorithms, such as BIdirectional Recurrent Neural networks with Adversarial Training (BIRNAT) (0.165 s) and Reversible SCI (RevSCI) (0.190 s). We believe that by fine-tuning the proposed network, we should be able to achieve SOTA results in video CS [79.80] and other reconstruction tasks [32.81-92].

During the review of our paper, we did notice that several new algorithms were proposed for spectral SCI reconstruction [35,93–96]. One of them used Transformer and brought competitive results to ours [93].

Regarding future work, advances in deep learning have empowered computational imaging for practical applications. Most recently, Transformer has shown promising performance

 Table 7. Extending Our Method for Video Compressive Sensing: Average PSNR, SSIM, and Running Time per

 Measurement of Different Algorithms on Six Benchmark Datasets

Algorithm	PSNR (dB)	SSIM	Running Time (s)
GAP-TV [7]	$26.73 \pm 4.33$	$0.858 \pm 0.082$	4.201 (CPU)
PnP-FFDNet [74]	$29.70 \pm 6.75$	$0.892 \pm 0.071$	3.010 (GPU)
DeSCI [8]	$32.65 \pm 7.07$	$0.935 \pm 0.047$	6180 (CPU)
BIRNAT [75]	$33.31 \pm 5.90$	$0.951 \pm 0.027$	0.165 (GPU)
U-net [76]	$29.45 \pm 4.75$	$0.882 \pm 0.057$	0.031 (GPU)
GAP-net-U-net-S12 [20]	$32.86 \pm 5.92$	$0.947 \pm 0.030$	0.03 (GPU)
MetaSCI [77]	$31.72 \pm 5.72$	$0.926 \pm 0.040$	0.025 (GPU)
RevSCI [78]	$33.92 \pm 6.02$	$0.956 \pm 0.025$	0.190 (GPU)
Ours	$33.53 \pm 5.90$	$0.954 \pm 0.026$	0.064 (GPU)



Fig. 8. Reconstructed frame of our method and other algorithms (GAP-TV, DeSCI, PnP-FFDNet, U-net, BIRNAT, RevSCI) on six benchmark datasets.

Table 8.Computational Complexity and AverageReconstruction Quality of Several SOTA Algorithms onSix Grayscale Benchmark Datasets

Algorithm	Params (10 <sup>6</sup> )	FLOPs (10 <sup>9</sup> )	PSNR (dB)	SSIM
BIRNAT [75]	4.13	390.56	33.31	0.951
U-net [76]	0.82	53.63	29.45	0.882
GAP-net-U-net-	5.62	87.58	32.86	0.947
S12 [20]				
MetaSCI [77]	2.89	54.16	31.72	0.926
RevSCI [78]	5.66	766.95	33.92	0.956
Ours	10.51	113.75	33.53	0.954

on many vision problems mainly because of its strong capability of extracting features. The self-attention mechanism in Transformer can capture global interactions between contexts and thus has advantages for global and local, muti-scale, spatial– temporal, or other features extraction that is difficult to realize by normal CNN-based networks. This can also inspire us to design new computational imaging systems. Specifically, the sampling process should be able to play the role of the first layer in Transformer to extract global or local features of the desired scene.

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